### Atomes et molécules

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## Université Paris Est Créteil, Créteil, France 2019-2020

# Chapter 2 : Evolution of atomic models

### Introduction

### 1 – State of knowledge at beginning of the 20th century

- 1.1 Rutherford model
- 1.2 Nature of light
- 1.3 Energy exchange matter/radiation
- 1.4 Emission spectrum of atomic hydrogen
- 1.5 Rutherford model problems

### 2 – Quantum model of the atom

2.1 Bohr model (1913)
2.2 Critics to the Bohr model
2.3 Bohr-Sommerfeld model
2.4 Critics to the Bohr-Sommerfeld model

### 3 – The electron in wave mechanics

3.1 Wave mechanics basics3.2 Wave model of the atom/Bohr quantum model3.3 Wave function definition

### Description of main steps of evolution of thoughts regarding the electronic structure of the atom ...



... example of the way in which science evolves.

Gathering of **experimental** data available at a given moment in time (observations/measurements).

Proposing a **theory**.

Construction of **models** (= mathematical formalization of the problem) allowing to interpret the whole or part of the experimental data.

A new model is necessary when the facts derived from new experiments cannot be explained by the model currently accepted.

A new good model never substitutes an old one: it includes it, completes it and refines it.

1 - State of knowledge ...

### 1.1 Rutherford model

Application of classical mechanics laws

Calculation of electron total energy ( $E_T = E_K + E_P$ )



$$E_{\text{total}} = -\frac{1}{8\pi\varepsilon_0} \frac{e^2}{r} = \frac{\cos\tan t}{r}$$

The electron energy is a continuous function of the distance electron – nucleus. According to electromagnetic laws , the electron should radiate, loose energy and fall on the nucleus !!! Another model is necessary

### Wave nature of light

At the end of the 19<sup>th</sup> century James Maxwell defines light as an electromagnetic wave traveling at a constant speed in vacuum. This speed is c ~  $300000 \text{ km/s} (3 \times 10^8 \text{ m.s}^{-1}.)$ 



He defines an electromagnetic wave as an oscillation of interrelated electric and magnetic fields

Light (and any other wave) is characterized by its frequency  $\nu$  (number of oscillations per second of the light wave, measured in Hz) or by its wavelength  $\lambda$  (measured in m.)

$$\lambda(m)\nu(s^{-1}) = c(m/s)$$

### 1 - State of knowledge ...

### 1.2 Nature of light

Visible light is nothing more than a small window in all the possible electromagnetic waves (which also include gamma rays, X-rays, ultraviolet, infrared and radio waves)



### **Corpuscular nature of light**

1905 : Albert Einstein  $\rightarrow$  suggest that luminous energy is someway 'granular' in nature.

This 'grain of energy' will be called photon in 1926. Then a new particle is born. This particle has no mass.

Energy of light:

$$E = h \, \nu = h \frac{c}{\lambda}$$



c: speed of light in vacuum (m s<sup>-1</sup>) =( $3.10^8$  m.s<sup>-1</sup> h: Planck's constant (Joule s) =  $6,62 \cdot 10^{-34}$  J.s  $\lambda$ : wavelength of associated radiation (m) 1 - State of knowledge ... - 1.3 Energy exchange matter/radiation

1900: Max Planck  $\rightarrow$  postulate the idea that energy exchange (absorption or emission) can be given only by an integer multiple of a minimal quantity of energy (quantum).

He wrote :

### $\Delta E = nh\nu$

 $\Delta E$ : variation in exchanged energy v: frequency of the emitted/absorbed radiation h: Planck's constant (6.62x10<sup>-34</sup> Joule.second) n: an integer number

This means that the radiative energy is quantified, i.e., that it cannot be exchanged in any arbitrary amount.

1 - State of knowledge... - 1.4 Emission spectrum atomic hydrogen

Experiments show that when gases are held at a small pressure in a sealed tube and subject to an appropriate electrical potential difference, they emit light.

Emission spectrum of atomic hydrogen



There are 4 lines in the visible domain. Their names: H $\alpha$ , H $\beta$ , H $\gamma$  and H $\delta$ , occurring at 656, 486, 434 and 410 nm respectively.

1 - State of knowledge... - 1.4 Emission spectrum: atomic hydrogen



Studying the spectrum over a wide spectral domain (UltraViolet, Visible, Infrared) shows that the wavelength of the emitted radiation satisfies the Ritz relation:

An empirical general law wich describes the whole spectrum 
$$\frac{1}{\lambda} = R_H \left(\frac{1}{p^2} - \frac{1}{q^2}\right) \qquad p, q \text{ integers, } p < q \\ RH: Rydberg \text{ constant} = 109677,76 \text{ cm}^{-1}$$
If  $p = 2$  and  $q = 3$   $\frac{1}{\lambda} = R_H \left(\frac{1}{3^2} - \frac{1}{2^2}\right) = 109677.76(1/4 - 1/9) = 15323$ 
 $1/15323 = 6.56 \ 10^{-5} \text{ cm} = 656 \ \text{nm}$  (H $\alpha$ )

1.5 Rutherford model problems



### Contradiction with electromagnetic theory

From Rutherford we know that the e- orbit the nucleus with uniform speed, describing a circular orbit.

It moves in an electromagnetic field

From electromagnetism theory, the e- must emit radiation with an associated loss of energy. This would imply a slow down of the movement and a decrease in the radius of the movement.

### ↓

The e- should 'fall' over the nucleus

Experience shows that, in the fundamental state, the atom does not radiate:

It is stable and its half life seems to be infinite

1 - State of knowledge... - 1.5 Emission spectrum atomic hydrogen



### Incoherence with the hydrogen emission spectrum

The excited H atom emits a radiation with energy  $E=hc/\lambda$  when an e- shifts from a level of energy to a lower one

$$\mathbf{E} = \mathbf{h}\mathbf{c}/\lambda = \mathbf{E}_{\text{total}(2)} - \mathbf{E}_{\text{total}(1)}$$
From Rutherford's model:  

$$E_{\text{total}(1)} = -\frac{1}{4\pi\varepsilon_0} \frac{e^2}{r_1} \quad \text{and} \quad E_{\text{total}(2)} = -\frac{1}{4\pi\varepsilon_0} \frac{e^2}{r_2}$$

Therefore:

$$\frac{1}{\lambda} = -\frac{e^2}{4\pi\varepsilon_0 hc} \left(\frac{1}{r_2} - \frac{1}{r_1}\right)$$

As there is no constraint over the values of the radii, the atomic hydrogen spectrum should be continuous!!!

### 2 – Quantum model of the atom -

2.1 Bohr model (1913)

1913: Niels Bohr  $\rightarrow$  proposes a model that avoids the 2 critics to the Rutherford model and presents the Bohr postulates:

\* The electron moves in a circular orbit around the nucleus, under the coulombic influence of the later, following classical mechanics

\* Only certain electronic orbits are allowed: those for which the angular moment of the electron L is a multiple of  $h/2\pi$ 

QUANTIFICATION CONDITION :  $L = mvr = nh/2\pi$ 



h: Planck constant (6.626x10<sup>-34</sup> J.s) n: integer m: mass v: velocity

\* In this allowed trajectory the electron does not emit radiation: it is on a stationary state (Etotal = constant)

Radiation is emitted/absorbed when an lectron moves discontinuously from a rate with energy Ei to one with energy Ef.  $v = \frac{E_f - E_i}{h}$  he frequency of the radiation is:

16

For the change of orbit to take place, the energy of the absorbed/emited photon has to be equal to the energy difference between the orbits



### **Bohr's model consequences**

The total energy of the electron is quantified (depends on an integer value n):

$E_{total} = -$	$me^4$ 1	cons tan te	13.6 <i>eV</i>	
	$-\overline{8arepsilon_0^2 h^2} \overline{n^2}$ –	$-\frac{1}{n^2}$	$-\frac{1}{n^2}$	

With this relation, Bohr explains the Ritz empirical relation:

$$\begin{array}{ccc}
 E_{i} & \longrightarrow & E_{i} - E_{f} = \frac{hc}{\lambda} = -\frac{me^{4}}{8\varepsilon_{0}^{2}h^{2}} \frac{1}{n_{i}^{2}} - \left(-\frac{me^{4}}{8\varepsilon_{0}^{2}h^{2}} \frac{1}{n_{f}^{2}}\right) \\
 E_{f} & & \frac{1}{\lambda} = -\frac{me^{4}}{8\varepsilon_{0}^{2}h^{2}} \left(\frac{1}{n_{f}^{2}} - \frac{1}{n_{i}^{2}}\right) \\
 Which is the same as & & \frac{1}{\lambda} = R_{H} \left(\frac{1}{p^{2}} - \frac{1}{q^{2}}\right) & \text{if} & R_{H} = -\frac{me^{4}}{8\varepsilon_{0}^{2}h^{2}}, n_{f} = p, n_{i} = q
\end{array}$$

Calculating  $R_H$  from this expression, Bohr finds a value in excellent agreement with the experimental value.

### 2 – Quantum model of the atom

Bohr's model predicts the atomic hydrogen spectrum



### 2 – Quantum model of the atom - 2.2 Critics to the Bohr model

Magnetic

field off.

Magnetic field on.



**n: principal quantum number** defining the smaller axis of the ellipse (b) *l*: orbital quantum number defining the flattening factor of the ellipse (1-b/a) **m: magnetic quantum number** defining the ellipse orientation with respect to the electric or magnetic field.

### Spin quantum number: s

### **Electron Intrinsic Angular Momentum**

Experimental evidence like the hydrogen fine structure and the Stern-Gerlach experiment suggest that an electron has an intrinsic angular momentum, independent of its orbital angular momentum. These experiments suggest just two possible states for this angular momentum, and following the pattern of quantized angular momentum, this requires an angular momentum quantum number of 1/2.

With this evidence, we say that the electron has spin  $\frac{1}{2}$ :

### Electron spin: s = -1/2 or s= +1/2 Spin "up" and "down"



### **Introduction of 4 quantum numbers**

Following the Bohr-Sommerfeld model the state of an electron in an atom is defined by 4 quantum numbers: **n**, *l*, **m and s**.

The values of these 4 numbers 'identify' an electron in an atom (as a citizen can be identified by an address: city, street, number, floor)

This result will not be contradicted nor modified by the wave model (that we will see next) which contributed to a new concept concerning the electronic organization of atoms.

### Principal quantum number: n

n can take all integer positive values. n = 1,2,3,... inf

**n** is the only quantum number that influences the energy of an electron in a hydrogen atom (Energy depends only on **n**)

**n** defines a 'shell' of electrons.

 $E_{total} = \frac{-13.6 \ eV}{n^2}$ 

A capital letter is assigned to each **shell** corresponding to a value of **n**:

n	1	2	3	4	5	6	•••
shell	K	L	Μ	Ν	0	Р	•••

2 – Quantum model of the atom - 2.3 Bohr-Sommerfeld model

### Secondary (orbital) quantum number: *l*

l can take all integer values between 0 and **n**-1.  $l = 0, 1, 2, 3, \dots n-2, n-1$ . Example: If n = 2, l = 0 or 1 l defines a sub-shell of electrons

A lowercase letter is assigned to each sub-shell corresponding to a value of  $\boldsymbol{\ell}$ 



Each sub-shell is named by a symbol composed of: \_ a number, the corresponding value of **n** 

\_ a letter, the corresponding value of  $\ell$ 

Exemple: sub-shell 1s (n = 1 et  $\ell$  = 0), sub-shell 3d (n = 3 et  $\ell$  = 2) Note: The electrons occupying these shells have the same name. E.g.: electrons 1s, electrons 3d.

### Magnetic quantum number: m

**m** can take  $(2^{*l}+1)$  integer values between **-** l and **+** l

$$m = -\ell, -\ell+1, ..., -1, 0, 1, ..., \ell-1, \ell$$

### *m* characterizes the energy sub-levels in a sub-shell of electrons

When an external magnetic field is applied, lines like the  $(n=3 \rightarrow n=2)$  transition of hydrogen split into multiple closely spaced lines. This is due to the separation of the n=3 and n=2 energy levels according to the number of *m* possible *values*.

For 
$$n=3$$
,  $\ell$  maximum = 2, m = -2,-1,0,1,2

For n = 2, l maximum = 1, m = -1,0,1



### 2 – Quantum model of ... - 2.4 Critics to Bohr-Sommerfeld model

Multi-electron atoms: the Bohr-Sommerfeld theory cannot explain the multielectron atoms spectra, starting from the Helium spectra.

It will lead to and be replaced by a new model provided by wave mechanics.



### Wave-particle duality

Light shows properties that can be interpreted as waves (continuous phenomenon) and at the same time as a set of particles (discrete phenomenon).

A monochromatic light ray that propagates as a wave

characterized by its wavelength  $\lambda$ 

can be described from a particle point of view

by photons of energy  $E = h \frac{c}{2}$ 

### 3 – The electron in wave mechanics - 3.1 Wave mechanics basics

**1924:** Louis de Broglie  $\rightarrow$  generalizes this idea of duality, suggesting that every material particle (also the electrons) can be considered to have wave and particle properties.

He relates the two descriptions by the relation:

$$\lambda = \frac{h}{p} = \frac{h}{mv}$$

$$\lambda$$
 : wavelength p: linear momentum





Every particle with linear momentum p is associated to a wave of wavelength  $\lambda$ 

Looking back to Bohr's condition on the angular momentum :  $L = mvr=nh/2\pi$  we can use  $\lambda = h/mv$  and we find  $2\pi r= n\lambda$ . The circumference of the electron orbit is a multiple of the electron wavelength. This is the stationary wave condition, a resonance phenomenon.

Wavelength for different states of the hydrogen atom:



### Heisenberg inequality (or uncertainty principle)

1927: Werner Heisenberg  $\rightarrow$  The momentum and position of a particle cannot be known simultaneously

$$\Delta x \Delta p > \frac{h}{2\pi}$$

 $\Delta x$ : position uncertainty  $\Delta p$ : momentum uncertainty



Applied to the electron this means we cannot know exactly where it is. We can only give the probability that it is at a certain point in space around the nucleus. The velocity is never zero!

### 3 – The electron... - 3.2 Wave model of the atom/Bohr quantum model

The idea of quantized electron energy is **conserved** (the energy can only take certain values, defining the energy levels among which the electrons are distributed)

BUT!

The idea of stable and geometrically well defined orbits for the electrons: \_ circular (Bohr) \_ elliptical (Bohr-Sommerfeld) is **abandoned**:

it is replaced by the concept of **probability of presence**.

### 3.3 Wave function definition



### The atom in wave mechanics

In wave mechanics the e- is not described as a point mass with an associated trajectory, it is considered as a **wave**.

Therefore, the electron doesn't follow the laws of classical mechanics (like in the Bohr model). It follows **wave mechanics** laws.

The electromagnetic wave associated to the electron is a stationary wave. Its amplitude at each point in space is time independent. The amplitude is given by a mathematical function

the wave function,  $\Psi$  (x,y,z)

### 3 – The electron...

 $\Psi(x,y,z)$  can be positive, negative, complex  $\Psi$  has no physical meaning

BUT!

 $\Psi^2$ : probability density of finding the electron in a particular position in space.

The probability dP of finding the electron in an infinitesimal volume dV centered at some point in space (x,y,z) is given by:

### $dP = | \Psi^2 | dV$

There is a normalization condition: the probability of finding the electron somewhere in space should be one:

```
\int_{\text{space}} \Psi^2 \, d\mathbf{V} = 1
```

### **QCM 6**

L'absorption d'un photon de fréquence convenable par l'atome d'hydrogène (Z = 1) peut entraîner :

- A. Un déplacement de l'électron vers un niveau d'énergie supérieure.
- **B.** L'ionisation de l'atome d'hydrogène.
- C. Aucune modification.
- **D.** Un déplacement de l'électron vers une trajectoire plus proche du noyau.
- **E.** La formation d'un isotope de l'hydrogène (le tritium caractérisé par Z = 1, A = 3, N = 2).

### QCM 7

Lorsqu'un électron revient d'un niveau excité au niveau fondamental, il y a :

A. Emission d'une radiation électromagnétique B. Absorption d'une radiation électromagnétique

#### **QCM 10**

Parmi les propositions suivantes quelle(s) est (sont) celle(s) qui est (sont) exacte(s) concernant le "grain d'énergie" appelé photon en 1926 par A. Einstein :

- A. Il s'agit d'une particule chargée positivement.
- **B.** Il s'agit d'une particule chargée négativement.
- C. Il peut transférer son énergie à de la matière.
- **D.** L'énergie E d'un photon est définie par la relation  $E = h\nu = hc/\lambda$ ,  $\nu$  définissant la fréquence et  $\lambda$  la longueur d'onde du rayonnement qui lui est associé, h étant la constante de Planck et c la vitesse de la lumière dans le vide.
- **E.** Aucune des propositions précédentes (de A à D) n'est exacte.

#### QCM 1

L'expression littérale générale des valeurs d'énergie permises pour l'électron de l'atome d'hydrogène dans le modèle de Bohr est du type :

$$\Box \quad E_n = +\frac{A}{n^2} \qquad \Box \quad E_n = +\frac{A}{2n} \qquad \Box \quad E_n = -\frac{A}{n^2} \qquad \Box \quad E_n = +\frac{A}{n} \qquad \Box \quad E_n = -\frac{A}{n^2}$$

On rappelle que A est une constante et n est le nombre quantique principal

#### QCM 2

Parmi les propositions concernant les nombres quantiques n (principal), l (secondaire) et m (magnétique), laquelle(lesquelles) est(sont) exacte(s)?

- $\Box$  *l* est toujours inférieur ou égal à *n*.
- $\square$  *m* peut prendre (2n+1) valeurs entières.
- $\Box\,$  Dans un atome, toutes les sous-couches s ont la même valeur de l.
- $\Box$  Un couple (n,l) donné définit une case quantique.
- $\Box$  Le couple (n = 2, l = 1) définit une case quantique 2p.