



Journal of Molecular Spectroscopy 245 (2007) 109-114

Journal of
MOLECULAR
SPECTROSCOPY

www.elsevier.com/locate/jms

Infrared band intensities of cyanobutadiyne (HC₅N) between 400 and 4000 cm⁻¹

Yves Bénilan a,*, Antoine Jolly a, Yann Trolez b, François Raulin a, Jean-Claude Guillemin b

^a LISA, Universités Paris 12 et Paris 7, CNRS, 61 Av. du Général de Gaulle, 94010 Créteil, France
 ^b Sciences Chimiques de Rennes, Ecole Nationale Supérieure de Chimie de Rennes, CNRS, 35700 Rennes, France

Received 3 July 2007; in revised form 19 July 2007 Available online 9 August 2007

Abstract

We report the first infrared spectra of pure gaseous cyanobutadiyne between 400 and 4000 cm^{-1} . A great care has been taken to obtain a pure sample so that, for the first time, intensities could be determined for all the main bands in the studied domain. The results are compared with available theoretical works. A sample containing a mixture of HC_5N and DC_5N was also studied in the same wavenumber range. The data obtained in this study should be very useful to identify those compounds by IR spectroscopy and will allow the quantification of cyanobutadyine in various environments. © 2007 Elsevier Inc. All rights reserved.

Keywords: HC₅N; DC₅N; Cyanobutadiyne; Deuterocyanobutadiyne; Infrared spectra; Integrated intensities

1. Introduction

Cyanobutadiyne (HC \equiv C-C \equiv C-CN, also called 2,4-pentadiynenitrile or cyanodiacetylene) was observed in space for the first time in the molecular cloud Sgr B2 by Avery et al. [1]. In the following years, its presence was acknowledged in numerous environments from dark molecular clouds to the envelop of dying carbon stars. Since its abundance is high enough to be observed in numerous environments, cyanobutadiyne is a good probe to compare the different physical conditions of those environments [2]. As an example, the recent study by Pardo et al. [3] analysed the rotational lines to deduce the physical and chemical structure of the envelop of the proto planetary nebula CRL618. In 1981, using rotational transitions, the deuterated specie (DC₅N) was observed in TMC 1 [4]. Its abundance compared to HC₅N is of the order of 1%.

In a totally different environment, results from simulation experiments of Titan's atmospheric chemistry also suggested the presence of HC₅N in the atmosphere of this

satellite of Saturn [5]. However, the attempts to detect it using millimeter observations were up-to-now unsuccessful [6].

The first spectroscopic study of cyanobutadiyne was reported by Alexander et al. [7], 30 years ago. In this study, they obtained the spectrum of a typical linear molecule in the ground state. From their data they derived bond lengths, rotational constants and a large permanent dipole moment. They also observed the presence of numerous satellite lines for each fundamental rotational transition corresponding to transitions from excited vibrational states. The strongest lines came from the v_{11} , the lowest bending mode [8].

In 1981, Yamada and Winnewisser [9] observed in the laboratory for the first time in the infrared a vibrational band of HC_5N : the fundamental CN stretching mode v_2 at 2256.124 cm⁻¹. The presence of very low-lying bending modes was also confirmed experimentally in this study [10]. Haas et al. [11] reported the infrared spectra of cyanobutadiyne between 500 and 4000 cm⁻¹ at low resolution. The product was obtained from a mixture of HC_3N and acetylene submitted to a dc discharge. The spectra of the distilled products showed large amount of hexatriyne

^{*} Corresponding author. Fax: +33145171564. E-mail address: benilan@lisa.univ-paris12.fr (Y. Bénilan).

Table 1
Positions and intensities of the major infrared absorption bands of HC₅N

Mode	Vibration	Position (cm ⁻¹)		Absolute intensities (cm ⁻² atm ⁻¹) at 295 K		
		This work	Theoretical results	This work		Theoretical results*
				Value	Error	
$v_1/v_3 + v_5$	C—H str.	(R) 3333 (P) 3322	3349 ^b	298	5.7	362 ^b
$2v_5$	overtone	(R) 2339 (P) 2330	2262 ^b	31.5	2.4	121 ^b
v_2	CN str.	(R) 2257 (P) 2248	2319 ^b	140	4.0	138 ^b
v_3	str.	(R) 2192.5 (P) 2182.5	2198 ^b	31.9	0.2	22.5 ^b
v_4	str.	Not detected	2067 ^b	_	_	0.33^{b}
$2v_7$	overtone.	(R) 1276 (P) 1267	_	90.7	1.8	_
v_5	str.	Not detected	1147 ^b	_	_	0.41 ^b
$v_9 + v_{10}$	combin.	(R) 718 (P) 709	_	16.7	0.6	_
v_7	bend.	642.2	650.5 ^d	245.1	6.4	177°/318°
v_6	str.	(R) 609 (P) 600	602 ^b	13.3	0.4	1.4 ^b
v_8	bend.	500.9	501 ^d	19.3	0.5	$7.4^{\rm b}/55^{\rm a}$
v ₉	bend.	461.1	462.9 ^d	1.6	0.3	2 ^b /0.8 ^a
v_{10}	bend.	Out of range	254 ^d	_	_	27 ^b /43 ^a
v_{11}	bend.	Out of range	106.8 ^d	_	_	$0.8^{\rm b}/0.4^{\rm a}$

a [12,20].

 (C_6H_2) that precluded the identification of all the main bands. Consequently, they focused their analysis on ν_7 , the most intense bending mode. In this study, they reported the presence of a lot of accompanying hot bands arising

from the presence of very low lying bending modes.

Parallel to this experimental approach, quantum molecular calculations have been performed. The first theoretical study of vibrational levels in HC₅N molecule was done with the main goal of determining the population of excited levels through infrared pumping [12]. From their ab initio calculations, they determined vibrational frequencies and band intensities for the normal modes v_1 to v_{11} . Unfortunately, since they were interested mainly in modelling the radiative transfer in relatively cool star envelops as IRC +10216, they reported the position and intensity only for the bending modes v_7 to v_{11} (see Table 1). Nevertheless, they showed that radiative excitation through the v_7 band alone is more effective than the collisional excitation because of the large intensity of this absorption band. Since then, numerous theoretical studies have been done and confirmed the low energy v_{10} and v_{11} and the intense v_7 bending mode [13,14 and references therein].

Finally, only two experimental studies of the infrared spectra of cyanobutadiyne in the gas phase have been reported [9,10] and only the positions of the v_2 and the v_7 bands have been firmly confirmed. Furthermore, starting from a mixture of compounds, none of the studies were able to give the absolute intensities of the infrared bands.

The recently reported preparation of cyanobutadiyne allows the isolation of a sample for the first time and opens the way for studies with a pure and thermodynamically stable sample [15]. In the present work, we report the first infrared spectra of pure gaseous cyanobutadiyne between 400 and $4000~{\rm cm}^{-1}$. A great care has been taken to obtain a pure sample so that, for the first time, intensities could be determined for all the main bands in the studied domain. The results are compared with available theoretical works. A sample containing a mixture of HC5N and DC5N was also studied in the same wavenumber range.

2. Experimental

Cyanobutadiyne was synthesised by the reaction of 1,3-butadiynyltri-n-butylstannane with p-toluenesulfonyle cyanide [15]. The pure sample was selectively trapped at $-70~^{\circ}\text{C}$ under vacuum (0.1 mbar) and kept under liquid nitrogen.

Deuterocyanobutadiyne has been prepared by addition of deuterated water (1 mL) and small amounts of potassium carbonate (10 mg) to a solution of cyanobutadiyne (30 mg) in dichloromethane (3 mL). After 10 min of stirring at room temperature, the aqueous phase was removed and the sequence was repeated two times. Purification was performed by vaporization of the organic phase in a vacuum line equipped with a tube ($L=15 \, \mathrm{cm}$, $\phi=30 \, \mathrm{mm}$) containing phosphorus pentaoxide in half-section and

^b [13].

c [21].

^{*} Converted from km/mol to cm⁻²atm⁻¹.

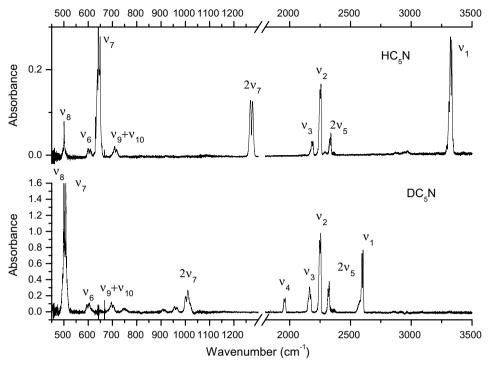


Fig. 1. Comparison between HC₅N and DC₅N spectra at 295 K and 0.5 cm⁻¹ resolution, between 450 and 3500 cm⁻¹.

followed by a selective condensation of deuter-ocyanobutadiyne in a $-70\,^{\circ}\text{C}$ cooled trap (for a similar experiment see [16]). The isotopic purity is of about 80%.
¹³C NMR (CD₂Cl₂, 293 K) δ 48.9 (C—CN), 65.6 (HC \equiv C, $^2J_{\text{CD}} = 7.8 \; Hz$), 67.2 (C \equiv C—CN), 72.1 (D—C, $^1J_{\text{CD}} = 40.5 \; \text{Hz}$), 105.2 (CN).

All spectra were recorded on an IRTF spectrometer (Bruker Equinox 55) equipped with a KBr beamsplitter and a DTGS detector. The spectrometer was flushed with nitrogen at a rate of 3 L/min in order to desaturate and stabilise water vapour and carbon dioxide absorption bands. The spectra were measured at room temperature and at 0.5 cm⁻¹ resolution between 400 and 4000 cm⁻¹. The gas samples were placed in a White cell (Specac 10.6 m pathlength) equipped with KBr windows. Sample pressures were measured with two MKS Baratron covering the 1 bar to 10⁻⁴ mbar pressure range. The gas samples were transferred from the tubes where they were stored in solid phase at liquid nitrogen temperature, to the cell via a vacuum line pumped below 10⁻⁷ mb with a turbo pump (Pfeiffer TMH 065).

The first infrared spectra of HC_5N showed the presence of some impurities in the sample. Butadiyne was easily identified and a distillation under vacuum at $-70\,^{\circ}C$ allowed to get rid of it. Another remaining impurity was not clearly identified. We suspected intermediate chlorine compound formed in the synthesis of the stannane precursor. Such compounds were already observed as residue in the synthesis of polyynes [17,18]. The purification procedure was performed by warming the sample at temperature around $-50\,^{\circ}C$ up to a pressure of 0.1 mbar in the 2.5 L volume of our apparatus. The spectrum was recorded

and checked to detect the presence of impurities. The sample was then pumped out and the procedure was repeated until the mixing ratio of the impurity was observed to be less than 0.04. This mixing ratio was checked afterward since the spectrum of the unidentified impurity can be obtained from the difference of highly polluted sample to purified ones.

For DC₅N, the sample purity was difficult to establish because of the limited amount of product. The sample showed the presence of both the chlorine residue and HC_5N . We grossly determined that 15% of HC_5N was present in the sample. We did not achieve the total purification of the sample and we estimated to 5% the amount of chlorine compound left in the sample. The DC₅N spectrum presented in Fig. 1 has been obtained after subtraction of the HC_5N spectrum and the unidentified impurity. No absolute intensities could be determined.

3. Spectral results

3.1. Bands position and relative intensities

 HC_5N is an unsymmetrical linear molecule with seven atoms. It has six fundamental stretching vibrational modes and five degenerate bending modes. The most recent theoretical determination of the frequency of those fundamental bands and some of their combination [13] indicates that the first two fundamental stretching modes (v_1 and v_2) are in fact dyads with the combination $v_3 + v_5$ and the overtone $2v_5$, respectively. Those are clearly identified on our spectra (Figs. 1, 2). As we will see later, the v_5 band was too weak to be observed. Thus, the observed $v_3 + v_5$

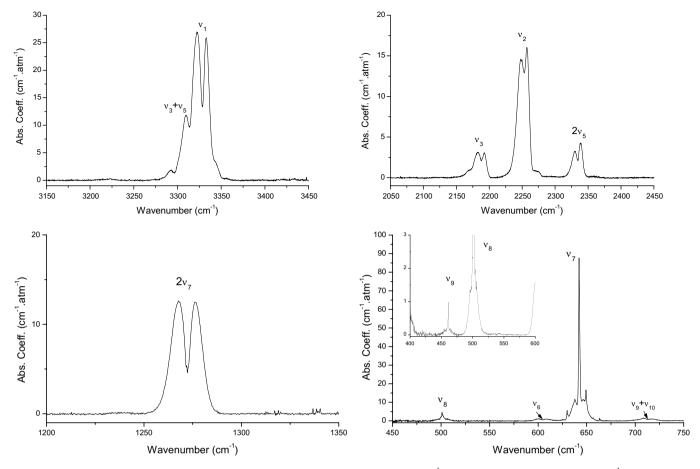


Fig. 2. Absolute absorption coefficient of HC₅N (base e at 295 K) at 0.5 cm⁻¹ resolution between 450 and 4000 cm⁻¹.

and $2v_5$ bands must be due to intensity borrowing through the Fermi resonance from the v_1 and v_2 bands, respectively.

The combination band $v_3 + v_5$ strongly overlaps the v_1 CH stretching mode so that the intensity is determined for the two bands together. Nevertheless, by comparing the maximum of the R branch of v_1 at 3333 cm⁻¹ and the maximum of the P branch of $v_3 + v_5$ at 3310 cm⁻¹ the fundamental stretching v_1 can be estimated to be at least two times more intense than the combination $v_3 + v_5$. In the DC₅N spectrum, since the v_1 stretching mode shifts to lower energies (2599 cm⁻¹), there is no Fermi resonance with the combination band which is not observed but predicted at higher energies. The next stretching vibration v_2 can be identified at 2253 cm⁻¹ for HC₅N and 2251 cm⁻¹ for DC₅N, as the lower component of the dyad. Its intensity is four times that of the overtone $2v_5$ observed at 2334 cm^{-1} for HC₅N and 2321 cm^{-1} for DC₅N. The v_3 stretching mode is also clearly observed as it already was in Hass et al. [11] low-resolution spectrum of HC₅N. The intensity of v_3 is approximately equal to that of the $2v_5$ overtone for both molecules.

The next two stretching modes could not be unambiguously identified in HC₅N. Their predicted intensities are at least ten times less than the previous ones [13]. So, they could be easily mixed up with combination bands or

absorption of the impurity. Nevertheless for DC₅N, the stretching vibration v_4 was predicted with intensity almost equal to that of stretching vibration v_3 [13]. It is observed in fact at 1960 cm⁻¹. The next band to be observed in HC₅N is the $2v_7$ overtone at 1272 cm^{-1} . It has an intensity ratio compared to the fundamental v₇ CCH bending mode which is about one third, approximately the same as the one observed between the equivalent v₅ fundamental CCH bending mode and its 2v₅ overtone for HC₃N [19]. In addition to the strong v_7 bending mode at 642 cm⁻¹, we also observe two more bending modes, v_8 at 501 cm⁻¹ and v_9 at 461 cm⁻¹ (Fig. 2). Their intensities are respectively one order and two orders of magnitude weaker than v_7 . We also observe the $2v_7$ overtone at 1005 cm⁻¹ for DC₅N. Its intensity is one third of the total intensity of the v_7 and v_8 bending modes, respectively, at 508 and 500 cm⁻¹. The proximity of those two modes leads to large interactions. Finally, two bands with low intensities and parallel band profiles are present both in HC₅N and DC₅N. The first one is observed around 700 cm⁻¹ and is assigned to a combination between the v_9 and the v_{10} modes. The second one is seen around 600 cm⁻¹ and is assigned to the fundamental v₆ stretching vibration. Finally, HC₅N has two more bending modes expected at 254 and 106.8 cm⁻¹ which are out of reach of our instrumental apparatus.

3.2. Absolute band intensities

All spectra correspond to the sum of two hundred scans. Eight different pressures, with column densities from 5×10^{-3} to 0.5 cm atm, have been used to determine the integrated intensities for the main bands. For the highest column densities used here, one bar of nitrogen was added to the sample to check for saturation effect. Nevertheless, no variation was observed. Band intensities $S_{\rm band}$ in cm⁻²atm⁻¹ were calculated as follows (ν is in cm⁻¹)

$$S_{\text{band}} = \frac{1}{pl} \int_{\text{band}} \ln \left(\frac{I_0}{I} \right) dv, \tag{1}$$

where I_0 is the incident intensity, I the transmitted intensity with the sample at the pressure p (in atm) and I is the pathlength (in cm). The results are reported in Table 1 for HC_5N . For each spectrum, the pressure is corrected from the presence of the impurity. This correction is only of a few percent since the purity of the sample varies from 96% to 98%. The mean error on the final intensities is determined from the $I\sigma$ standard deviation on the data. It is generally less than 3% except for the $2v_5$ overtone, which is overlapped by CO_2 absorption, and for the v_9 , which has a very low intensity. Since for DC_5N , we only have one sample and no possibility to estimate the purity, we report in Table 2 the relative intensities integrated for each band after subtraction of the HC_5N spectrum.

3.3. Comparison with theoretical studies reported in the literature

The most recent theoretical studies used large-scale coupled cluster calculation of CCSD(T) type with cc-pVTZ basis set [13] or cc-pVQZ [14], in order to predict harmonic

vibrational wavenumbers. Anharmonic perturbations were predicted by Botschwina et al. [13] for the two dyads $(v_1,$ $v_3 + v_5$) and $(v_2, 2v_5)$ and for the v_3 stretching mode. The measured positions of the bands are generally in good agreement with the anharmonic calculations. However, systematic differences persist between theoretical and experimental positions. They are about 20 cm^{-1} for v_1 and $v_3 + v_5$, and 10 cm^{-1} for v_2 and v_3 . The observed band at 2335 cm⁻¹ assigned to $2v_5$ shows a shift of -15 cm⁻¹ which can be explained by strong anharmonic interaction with v_2 . The use of cc-pVQZ basis set [14] does not change much the conclusions since the calculated harmonic frequency for the v_1 to v_3 stretching mode are not much different. Concerning the intensities, ab initio calculations are less efficient to predict accurate values. The intensities of the two dyads are overestimated by 30% and the intensity of v_3 is underestimated by about the same amount. The same authors have also calculated the harmonic positions of the bending modes. Here again, the positions are efficiently estimated. Since the anharmonicity decreases for the low energy modes, harmonic calculations using ccpVQZ basis set [14], are able to predict the position of v_8 and v_9 with a precision of 1 or 2 cm⁻¹. Consequently, the position of v_{10} and v_{11} , which we could not verify, calculated by the same authors at respectively 254 and 106.8 cm⁻¹, are certainly accurate values. Intensities of the five bending modes were determined by Uvemura [20] using semi empirical calculations and more recently by Scemama et al. [21] using B3LYP/cc-pVQZ method. The first study overestimates the intensity of v_7 by 30% whereas the second underestimates it by the same amount. The errors are even greater when looking at v₈ and v₉. Nevertheless, one can notice that the results given by those theoretical methods always bracket the measured value. One

Table 2 Positions and relative intensities of the major infrared absorption bands of DC₅N

Mode	Vibration	Position (cm ⁻¹))	Relative intensities		
		This work	Theoretical results	This work (Uncertainty = 10%)	Theoretical results ^c	
v_1	C—D str.	(R) 2603 (P) 2595	2617 ^b	0.86	0.66 (124.5) ^b	
$2v_5$	overtone	(R) 2325 (P) 2317	2251 ^b	0.30	0.40 (75) ^b	
v_2	CN str.	(R) 2255 (P) 2246	2308 ^b	1	1 (188.5) ^b	
v_3	str.	(R) 2173.5	2173 ^b	0.35	0.23 (43.5) ^b	
v_4	str.	(P) 2164 (R) 1964 (P) 1955	1964 ^b	0.17	0.18 (34) ^b	
$2v_7$	overtone	(R) 1010 (P) 1001	_	0.30	_	
v_6	str.	(R) 604 (P) 595	595 ^b	0.11	0.006 (1.1) ^b	
v_7	bend.	508	516.9 ^b	1.05	0.94 (168.5) ^a	
v_8	bend.	500	505.1 ^b	1.05	$0.94 (9.5^{a})$	

a [12,20].

b [13]

^c Relative intensities (absolute intensities converted from km/mol to cm⁻²atm⁻¹).

would then predict that the intensity of the v_{10} and v_{11} mode should be respectively around 35 and 0.6 cm⁻² atm⁻¹ with an uncertainty of the order of 30%.

4. Conclusion

In this experimental study, we have obtained for the first time the spectra of isolated cyanobutadiyne and its deuterated isotopomer (DC₅N), in the 400 to 4000 cm⁻¹ wavenumber range. The preparation of the HC₅N sample allowed obtaining purity better than 97%. This opens the way to the determination of the position and the intensity of the fundamental vibrational modes present in the studied wavenumber range. The v_1 , v_2 , v_3 , and v_6 stretching modes could be identified and their intensities determined. The missing modes v_4 and v_5 have too low intensities to be observed. The v_7 , v_8 , and v_9 bending modes were studied and their intensities determined. Beside those fundamental modes, we have also identified two overtones, namely $2v_5$ and $2v_7$, and two combination bands: $v_3 + v_5$ and $v_9 + v_{10}$. The v_{10} and v_{11} were out of the energy range studied here. For DC₅N we obtained a sample with purity better than 80%. This allowed us to subtract the contribution of HC₅N from the spectra and thus to identify most of the fundamental modes present in the studied wavenumber range. Only the v_5 mode is missing. Nevertheless, we observed the two same overtones and combination bands as in HC₅N.

The comparison of our experimental study with previous theoretical predictions lead to the conclusion that for such molecule, high level *ab initio* method like couple cluster calculations are able to predict band positions with an error less than 1%. On the other hand, we found differences as large as 30% between experimental determinations and theoretical predictions concerning the band intensities.

The data obtained in this study should be very useful to identify HC_5N by IR spectroscopy. As an example, a recent study of Coupeaud et al. [22] already used those data to confirm their identification of HC_5N in the irradiation of C_2H_2/HC_3N or C_2H_2/C_4N_2 mixture in cryogenic matrix. The determination of band intensities also opens the way to the quantification of this compound. As an example, Cernicharo et al. [23] used the intensity values of Deguchi and Uyemura [12] to determine the abundance of HC_5N in the spectra of CRL618 proto planetary nebula obtained with ISO. Since the values derived by Deguchi and Uyeruma [12] for the ν_7 are 30% too high, the derived column densities are underestimated by the same amount. This could have some implication for the modelled photochemistry in this environment.

Another environment in which cyanobutadyine might be observed using infrared spectroscopy is Titan's atmosphere where an active photochemistry of nitrogen and methane is producing complex molecules. In fact, HC_5N has been observed in simulation experiments of such environment [5]. Even if the attempts to detect it using millimetre observations were unsuccessful [6], the Cassini-Huygens mission and its CIRS (Composite InfraRed Spectrometer) instrument are presently accumulating data in which the signature of HC_5N might raise from the noise level in a near future.

Acknowledgment

This work was financially supported by the CNES (Centre National d'Etudes Spatiales).

References

- [1] L.W. Avery, N.W. Broten, J.M. MacLeod, T. Oka, H.W. Kroto, Astrophys. J. 205 (1976) L173–L175.
- [2] D. Jenning, K. Fox, Astrophys. J. 254 (1982) 111-115.
- [3] J.R. Pardo, J. Cernicharo, J.R. Goicoechea, Astrophys. J. 628 (2005) 275–282.
- [4] J.M. MacLeod, L.W. Avery, N.W. Broten, Astrophys. J. 251 (1981) L33–L36.
- [5] E. de Vansay, M.C. Gazeau, J.C. Guillemin, F. Raulin, Planet. Space Sci. 43 (1995) 25–31.
- [6] A. Marten, T. Hidayat, Y. Biraud, R. Moreno, Icarus 158 (2) (2002) 532–544.
- [7] A.J. Alexander, H.W. Kroto, D. Walton, J. Mol. Spectrosc. 62 (1976) 175–180.
- [8] M. Hutchinson, H.W. Kroto, D.R.M. Walton, J. Mol. Spectrosc. 82 (1980) 394-410.
- [9] K. Yamada, G. Winnewisser, Z. Naturforsch. 36a (1981) 1052-1056.
- [10] K. Yamada, G. Winnewisser, Z. Naturforsch. 40a (1985) 476-478.
- [11] S. Haas, G. Winnewisser, K.M.T. Yamada, Can. J. Phys. 72 (1994) 1165–1178.
- [12] S. Deguchi, M. Uyemura, Astrophys. J. 285 (1984) 153-157.
- [13] P. Botschwina, A. Heyl, M. Oswald, T. Hirano, Spectrochim. Acta Part A 53 (1997) 1386–1425.
- [14] C. Degli Esposti, L. Bizzocchi, P. Botschwina, K.M.T. Yamada, G. Winnewisser, S. Thorwirth, P. Forster, J. Mol. Spectrosc. 230 (2005) 185–195
- [15] Y. Trolez, J.-C. Guillemin, Angew. Chem. Int. Ed. 44 (2005) 7224–7226
- [16] Y. Bénilan, A. Jolly, F. Raulin, J.-C. Guillemin, Planet. Space Sci. 54 (2006) 635–640.
- [17] N. Smith, Y. Bénilan, P. Bruston, Planet. Space Sci. 46 (1998) 1215– 1220.
- [18] F. Shindo, Y. Bénilan, J.-C. Guillemin, P. Chaquin, A. Jolly, F. Raulin, Planet. Space Sci. 51 (2003) 9–17.
- [19] A. Jolly, Y. Benilan, A. Fayt, J. Mol. Spectrosc. 242 (2007) 46-54.
- [20] M. Uyemura, Bull. Chem. Soc. Jpn. 59 (1986) 1303-1309.
- [21] A. Scemama, P. Chaquin, M-C. Gazeau, Y. Bénilan, J. Phys. Chem. A 106 (15) (2002) 3828–3837.
- [22] A. Coupeaud, R. Kolos, I. Couturier-Tamburelli, J.-P. Aycard, N. Piétri, J. Chem. Phys. A 110 (2006) 2371–2377.
- [23] J. Cernicharo, A.M. Heras, A.G. Tielens, J.R. Pardo, F. Herpin, M. Guélin, L.B. Waters, Astrophys. J. 546 (2001) L123–L126.